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# Gluino mass limits with sbottom NLSP in coannihilation scenarios

### Arindam Chatterjee,<sup>a</sup> Arghya Choudhury,<sup>a</sup> Amitava Datta<sup>b</sup> and Biswarup Mukhopadhyaya<sup>a</sup>

*E-mail:* arindam@hri.res.in, arghyachoudhury@hri.res.in, adatta@iiserkol.ac.in, biswarup@hri.res.in

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<sup>&</sup>lt;sup>a</sup> Regional Centre for Accelerator-based Particle Physics, Harish-Chandra Research Institute, Allahabad 211019, India

<sup>&</sup>lt;sup>b</sup>Department of Physics, University of Calcutta, 92 A.P.C. Road, Kolkata 700 009, India

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Introduction

#### Introduction 1

Supersymmetry (SUSY) [1–5] is the most popular and widely studied extension of the standard model (SM) of particle physics. In the R-parity conserving minimal supersymmetric standard model (MSSM), the lightest neutralino is the lightest SUSY particle (LSP) and a viable dark matter (DM) [6-11] candidate. R-parity makes the LSP stable leading to missing energy  $(\not\!\!E_T)$  signals at the Large Hadron Collider (LHC). Right from the beginning of the LHC run both the ATLAS and CMS collaborations are looking for SUSY using this feature of the signals in various channels. In the absence of any excess events they have put stringent bounds on the masses of supersymmetric particles (sparticles) from the 7 and 8 TeV data. The limits on the masses of strongly interacting sparticles are the strongest, due to the to large production cross-sections [12-19]. For example, at the end of the 8 TeV run,  $m_{squark}$   $(m_{\tilde{q}}) = m_{qluino}$   $(m_{\tilde{q}})$  scenarios are ruled out up to 1.7 TeV [12].

It should be borne in mind that comprehensive SUSY search strategies in the upcoming experiments at LHC 13 TeV runs will be designed on the basis of the exclusions obtained by the experiments during the first phase. It is, therefore, worthwhile to revisit the limits critically and find out the models in which some loopholes in the current search techniques significantly relax the limits. A case in point is the compressed SUSY scenario which can considerably relax the limits [20-26].

In view of the stringent bounds on the first two generation squark masses, models with light third generation have received much attention in recent times [27-51]. It may be recalled that such heavy squarks offer a way of ameliorating the SUSY flavour [52] and CP problems [53, 54]. Such scenarios also help in restoring the naturalness of a SUSY model.

In this analysis, we focus on scenarios with a light  $b_1$ , both with and without a  $\tilde{t}_1$ , along with a heavier gluino which exclusively decays into these squarks. We further assume that the remaining members of the MSSM spectrum are heavy, except the lightest neutralino which lies below the  $b_1$ . The  $b_1$  is the next-to-lightest SUSY particle (NLSP) in this scenario.

An added attraction, and at the same time a crucial viability check of this spectrum is that LSP–NLSP coannihilation can produce the observed DM relic density of the universe

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provided their mass difference is small. A large number of phenomenological analyses have already addressed various topics/issues related to stop searches in different channels [27– 42] and investigated DM production via  $\tilde{t}_1 - \tilde{\chi}_1^0$  coannihilation [37–42] extensively in the context of LHC. On the other hand the sbottom NLSP scenario, which also has the potential of explaining the observed relic density, has not yet received due attention and has been addressed by a relatively small number of analyses only [43-51]. Moreover, most of the analyses predating the first phase of the LHC experiments were restricted to mSUGRA motivated models with non-universal boundary conditions. Since the SUSY breaking mechanism is essentially unknown, our emphasis will be on the  $b_1 - \tilde{\chi}_1^0$  coannihilation scenario in the phenomenological minimal supersymmetric standard model (pMSSM) [55] constrained by the LHC data. Thus, clear perspectives are expected to emerge from this analysis on (a) the viability of sbottom-neutralino coannihilation (with or without the assistance of stops) from the standpoint of relic density, (b) the spectacular signatures, not viable in the  $t_1$ -NLSP scenario, expected in the phase 2 of the LHC experiments and c) the new constraints in the  $b_1 - \tilde{g}$  sector in the  $b_1 - \tilde{\chi}_1^0$  co-annihilation scenarios using the current LHC data.

It should be noted in addition that one of the scenarios investigated here, namely the one with just a (right)  $\tilde{b}_1$  and both stop mass eigenstates heavy, has an appeal from the viewpoint of the lighter neutral Higgs mass. It is known that pushing this mass upto  $\approx$  125 GeV becomes less troublesome if the stop mass(es) and the trilinear SUSY breaking term is large. It therefore helps to have a situation where both of the stop eigenstates participate in raising the Higgs mass, while a light  $\tilde{b}_1$  lies close to the lightest neutralino and and facilitates co-annihilation.

Before delving into our main analysis it is worthwhile to review the existing LHC limits on  $m_{\tilde{t}_1}$  and  $m_{\tilde{b}_1}$  from direct search of these sparticles. If  $\Delta m_{\tilde{t}_1} = m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0}$  is relatively large so that the decay  $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$  decay is allowed, stop masses in the range 210–650 GeV are excluded for  $m_{\tilde{\chi}_1^0} \lesssim 30 \text{ GeV}$  [56, 57]. As  $m_{\tilde{\chi}_1^0}$  increases the limits become weaker. Similarly from sbottom pair production  $m_{\tilde{b}_1} < 620 \text{ GeV}$  is excluded at 95% CL for  $m_{\tilde{\chi}_1^0} \lesssim 150 \text{ GeV}$  [58]. However, for  $m_{\tilde{\chi}_1^0} \gtrsim 250 \text{ GeV}$  there is no limit on  $m_{\tilde{t}_1}$  and  $m_{\tilde{b}_1}$ .

None of the above scenarios, however, is consistent with the observed DM relic density of the universe. To explain the correct relic density, we need small mass difference between LSP and the NLSP (stop/sbottom) and a different technique for the NLSP search is called for. The limits on masses on third generation squarks are remarkably weaker in such coannihilating scenarios. From direct stop pair production the latest bound on  $m_{\tilde{t}_1}$  is 240 GeV [59] for  $m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} < 80$  GeV. If the stop and the lightest neutralino are almost degenerate then stop masses up to 260 GeV are excluded from ATLAS search with a 'monojet like' topology<sup>1</sup> [59]. Very recently using this search channel and 20 fb<sup>-1</sup> data at 8 TeV,  $m_{\tilde{b}_1}$  below 250 GeV is excluded for small  $\Delta m_{\tilde{b}_1} = m_{\tilde{b}_1} - m_{\tilde{\chi}_1^0}$  [59].

As the gluino pair production cross-sections is the largest, gluino decays into the third generation squarks are likely to probe larger ranges of  $m_{\tilde{t}_1}$  and/or  $m_{\tilde{b}_1}$  in the coannihilation

<sup>&</sup>lt;sup>1</sup>Here one depends on a hard ISR jet in enhancing the signal. However, the signal with low jet multiplicity may contain more than one jet.

scenario. Many groups have looked for gluino decay signatures in presence of light third generation squarks [60–68]. In such cases gluino decays into  $\tilde{t}_1 t$  and/or  $\tilde{b}_1 b$ . For coannihilating scenarios  $\tilde{t}_1$  decays into  $c\tilde{\chi}_1^0$  and  $\tilde{b}_1$  decays into  $b\tilde{\chi}_1^0$  with 100% BRs. However the c or b jets coming from  $\tilde{t}_1$  or  $\tilde{b}_1$  will be softer and the limits on the gluino masses could be degraded. For example, if all three generations squarks are much heavier than the gluino, then BR( $\tilde{g} \to q\bar{q}\tilde{\chi}_1^0$ ) is 100% where q is any quark. In this case gluino masses below 1.4 TeV is excluded for  $m_{\tilde{\chi}_1^0} \leq 200 \text{ GeV}$  from  $jets + \not{E}_T$  channel [12]. On the other hand assuming  $m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} = 20 \text{ GeV}$  which is relevant for the  $\tilde{t}_1$  - LSP coannihilation and Br( $\tilde{g} \to \tilde{t}_1 t \to ct\tilde{\chi}_1^0$ ) = 100%, the limit reduces to 1150 GeV [12]. The ATLAS collaboration have also updated their search for NLSP shottom scenario in the  $0 - 1l + 3b - jets + \not{E}_T$  channel [13] with Br( $\tilde{g} \to \tilde{b}_1 b \to bb\tilde{\chi}_1^0$ ) = 100% and have excluded  $m_{\tilde{g}}$  below 1.2 TeV upto  $m_{\tilde{b}_1} = 1.0 \text{ TeV}$ . But the model is in conflict with the DM relic density as the LSP mass is fixed at 60 GeV.

In this paper we have revisited the last analysis focusing on  $\Delta m_{\tilde{b}_1}$  appropriate for the coannihilation scenario, and checked whether this relaxes the gluino mass limit. We have considered several variants of the sbottom NLSP scenario characterized by the presence or absence of a  $\tilde{t}_1$ . In contrast, the experimental searches have considered gluino decays either into the  $\tilde{t}_1$  or the  $\tilde{b}_1$  channel. But if both  $\tilde{t}_1$  and  $\tilde{b}_1$  are light, then, depending on their mass difference, the gluino may decay into both the channels with sizable BRs. If, for example, the stop exclusively decays into a soft charm jet the gluino mass limit may suffer further suppression. Moreover the presence of an additional light  $\tilde{t}_1$  predicts novel signal for the LHC phase 2 experiments as we shall elaborate below.

In section 2, we discuss the variants of the scenario studied here, and identify the regions of the parameter space which are consistent with the observed relic density as well as direct DM search experiments. We also identify the spectacular signatures in the  $\tilde{b}_1$ -NLSP scenario in presence of a relatively light  $\tilde{t}_1$ . The new constraints in the  $m_{\tilde{b}_1} - m_{\tilde{g}}$  plane in different scenarios using the LHC Run I data are computed in section 3. We summarize our main conclusions in section 4.

## 2 Co-annihilating $\tilde{b}_1$ and $\tilde{\chi}_1^0$ relic abundance: with and without light stop

We explore a (simplified) pMSSM scenario with a bino-like LSP  $(\tilde{\chi}_1^0)$  and  $b_1$  as NLSP. Since the bino couples to the other (s)particles only through U(1)<sub>y</sub> gauge interaction, the annihilation cross-section into standard model particles is known to be small. As a consequence, a (pure) bino-like  $\tilde{\chi}_1^0$  decouples from the thermal soup sufficiently early, and therefore has a larger thermal relic abundance compared to the present measured relic abundance  $\Omega h^2 = 0.112$  [69] over large region of the parameter space. However, when there is a scope for co-annihilation with the NLSP the DM and the coannihilating sparticle ( $\tilde{b}_1$  in this case) remain in relative thermal equilibrium for a longer period of time through  $\tilde{\chi}_1^0 SM \leftrightarrow \tilde{b}_1 SM'$ , where SM and SM' denote two Standard Model particles which are assumed to be in thermal equilibrium and therefore abundant. In the context of a bino-like  $\tilde{\chi}_1^0$ , the *s*-channel  $\tilde{b}_1$  exchange process  $\tilde{\chi}_1^0 b \leftrightarrow \tilde{b}_1 g$ , and the t/uchannel *b* exchange processes  $\tilde{\chi}_1^0 g \leftrightarrow b\tilde{b}_1$  are examples of this. Thus, larger (co-)annihilation cross-section of  $\tilde{b}_1$  also implies a lower thermal relic density of the  $\tilde{\chi}_1^0$ . Of course, these interactions eventually decouple and the remaining sbottoms decay to  $\tilde{\chi}_1^0$ . The effect can be captured by [70, 71] using an effective (co-)annihilation cross-section ( $\sigma_{\text{eff}}$ ),

$$\sigma_{\rm eff} = \sum_{i,j} \frac{g_i g_j}{g_{\rm eff}^2} (1 + \Delta_i)^{3/2} (1 + \Delta_j)^{3/2} e^{-x(\Delta_i + \Delta_j)} \sigma_{ij}, \qquad (2.1)$$

instead of the annihilation cross-section  $\sigma_{ann}$  in the relevant expressions. In the above equation,  $\{i, j\}$  runs over the list of co-annihilating sparticles,  $g_i$  denotes the number of degrees of freedom of the *i*-th sparticle,  $\Delta_i = \frac{m_i}{m_{\tilde{\chi}_1^0}} - 1$ ,  $x = \frac{m_{\tilde{\chi}_1^0}}{T}$  and  $\sigma_{ij}$  denotes the co-annihilation cross-section of *i* and *j*-th sparticles into SM particles. Also,

$$g_{\text{eff}} = \Sigma_i g_i (1 + \Delta_i)^{3/2} e^{-x\Delta_i}.$$
(2.2)

Thus co-annihilations are very relevant for sufficiently small  $\Delta_i$ , i.e. for small mass difference between the LSP and the NLSP. In this paper we choose the mass difference between  $\tilde{\chi}_1^0$ and  $\tilde{b}_1$  such that sufficient co-annihilation takes place [70, 71] ensuring the correct thermal relic abundance of  $\tilde{\chi}_1^0$ . Note that, LSP co-annihilation with strongly interacting sparticles the involves one strong vertex leading to large cross-section. Consequently to achieve the required  $\sigma_{\text{eff}}$ , the mass difference of these strongly interacting sparticles with the LSP need to be larger compared to their electroweak counterparts in a co-annihilation scenario. As we will see, while the required mass difference is about 5–15 GeV for electroweak sparticles (see, e.g., [72, 73]), for (third generation) squarks it is about 25–35 GeV. This has nontrivial consequences for the observability of  $\tilde{b}_1$  coannihilation scenario at the LHC (to be discussed in the next section).

In the pMSSM, the soft-supersymmetry-breaking mass terms for the third generation squarks, namely  $m_{\tilde{b}_R}$   $(m_{\tilde{t}_R})$  for the SU(2)<sub>L</sub> singlets  $\tilde{b}_R$   $(\tilde{t}_R)$  and  $m_{\tilde{Q}_3}$  for the SU(2)<sub>L</sub> doublets are free parameters. A large hierarchy among the parameters  $m_{\tilde{b}_R}$   $(\tilde{t}_R)$  and  $m_{\tilde{Q}_3}$  ensure that L-R mixing in the sbottom (stop) sector is small. We shall restrict ourselves to such scenarios only. We further assume that the first two generations of squarks are much heavier and decoupled. The electroweak sparticles other than the LSP are assumed to be heavier than  $\tilde{b}_1$  or  $\tilde{t}_1$ . We consider three  $\tilde{b}_1$ -LSP coannihilation scenarios characterized by different relative magnitudes of the above soft terms leading to potentially distinctive LHC signatures. Each scenario is illustrated by a BM point in table 1 and the corresponding BRs relevant for LHC signatures are tabulated in table 2.

• Gluino-Sbottom-R-Model: Here the  $\tilde{b}_1$  NLSP is a SU(2)<sub>L</sub> singlet with a small  $m_{\tilde{b}_R}$ . The parameters  $m_{\tilde{Q}_3}$  and  $m_{\tilde{t}_R}$  are assumed to be too large to affect the gluino decays. Benchmark point (BMP) BMP-A in table 1 provides an example of this scenario.

Parameter	BMP-A	BMP-B1	BMP-B2	BMP-B3	BMP-C
$m_{\widetilde{\chi}^0_1}$	300	306	305	305	305
$m_{\widetilde{q}}$	1236	1259	1273	1270	1310
$m_{\widetilde{b}_1}$	325	335	333	333	334
$m_{\widetilde{t}_1}$	1558	345	507	468	455
$\Omega_h^2$	0.11	0.11	0.11	0.11	0.12

**Table 1**: Mass spectra for different benchmark points. BMP-A,B(1,2,3),C represent three different scenarios Gluino-Sbottom-R-Model, Gluino-Stop-Sbottom-R-Model and Gluino-Stop-Sbottom-L-Model respectively (see text for details). All the masses are in GeV.

• Gluino-Stop-Sbottom-R-Model: In this model the NLSP is dominantly  $\tilde{b}_R$  as in the previous case. However, there is also a R-type light  $\tilde{t}_1$ . In contrast to the previous scenario,  $\tilde{g}$  decays to both  $b\tilde{b}_1$  and  $t\tilde{t}_1$  channels with sizable BRs. Depending on the mass splitting between  $\tilde{b}_1$  and  $\tilde{t}_1$ , there are three possible decay modes of  $\tilde{t}_1$ :  $c\tilde{\chi}_1^0$ ,  $t\tilde{\chi}_1^0$  and  $\tilde{b}_1W$ . They are tabulated as BMP-B1, BMP-B2 and BMP-B3 respectively in table 1.

If the mass difference between  $\tilde{b}_1$  and  $\tilde{t}_1$  is assumed to be significantly smaller than  $m_W, \tilde{t}_1$  decays dominantly into the loop induced mode  $c \tilde{\chi}_1^0$ . As a result the number of taggable hard jets in the signal decreases and this scenario is expected to yield the weakest limits on  $m_{\tilde{a}}$ . Moreover, both  $\tilde{t}_1$  and  $\tilde{b}_1$  can contribute to DM production if their mass difference is 10 GeV or so. BMP-B1 in the table illustrates this case. However, if  $m_{\tilde{t}_1} \stackrel{>}{\sim} m_t + m_{\tilde{\chi}_1^0}$ , then  $\tilde{t}_1 \to t \tilde{\chi}_1^0$  dominates; BMP-B2 illustrates this scenario. Finally, for  $m_W + m_{\tilde{b}_1} \stackrel{<}{{}_\sim} m_{\tilde{t}_1} \stackrel{<}{{}_\sim} m_t + m_{\tilde{\chi}_1^0}, \ \tilde{t}_1 \rightarrow \tilde{b}_1 W$  dominates over the flavor-violating decay mode  $c_{\tilde{\chi}_1^0}^{\gamma_1^0}$  provided there is a tiny L-R mixing in the stop and sbottom sectors. This is depicted by BMP-B3. Note that in this case the L-components in  $\tilde{t}_1$  and  $b_1$  are respectively enhanced to about 5% and 0.2% only by adjusting the parameters in the stop and sbottom mass matrices. However, if both the mode  $\tilde{t}_1 \to t \tilde{\chi}_1^0$  and  $\tilde{t}_1 \to b_1 W$  are kinematically allowed the mixing angle suppressed latter mode is not competitive. In the presence of large mixing in the sbottom and/or stop sector the above classifications become somewhat blurred. In this case, e.g, both  $\tilde{t}_1 \to b_1 W$  and  $\tilde{t}_1 \to t \tilde{\chi}_1^0$  may be observed, if kinematically allowed. In fact the relative rates of these two events may provide some information on L-R mixing in the squark sector.

• Gluino-Stop-Sbottom-L-Model: In this case only  $m_{\tilde{Q}_3}$  is relatively light. The  $\tilde{b}_1$  NLSP is dominantly an SU(2)<sub>L</sub> doublet i.e, of L-type. In the limit of small L - R mixing,  $\tilde{t}_1$  is heavier than  $\tilde{b}_1$  due to larger contributions from the D-terms and F-terms. The mass difference between  $\tilde{t}_1$  and  $\tilde{b}_1$  is fixed and only allows the decay  $\tilde{t}_1 \to \tilde{b}_1 W^+$  which occurs with 100 %. The BMP-C in table 1 illustrates this case.

The BRs relevant for LHC signatures are tabulated in table 2. In all scenarios,  $Br(b_1 \rightarrow b_2)$ 

Decay Modes	BMP-A	BMP-B1	BMP-B2	BMP-B3	BMP-C
$\widetilde{g} \to b\widetilde{b}_1$	100	52.3	52.2	50.8	54.4
$\rightarrow t \widetilde{t}_1$	-	47.7	47.8	49.2	45.6
$\widetilde{b}_1 \to b \widetilde{\chi}_1^0$	100	100	100	100	100
$\widetilde{t}_1 \to c \widetilde{\chi}_1^0$	-	97	-	-	-
$\rightarrow W\widetilde{b}_1$	-	3*	1.4	100	100
$ ightarrow t \widetilde{\chi}_1^0$	-	-	98.6	-	_

**Table 2**: Branching Ratios - BRs (%) of decay modes of  $\tilde{g}$ ,  $\tilde{b}_1$  and  $\tilde{t}_1$  for benchmark points. \* For BMP-B1 small BRs of  $\tilde{t}_1 \to W\tilde{b}_1$  is coming from virtual W.

 $b\tilde{\chi}_1^0$ ) is 100%, as  $\tilde{b}_1$  is the NLSP. In the scenarios with a light  $\tilde{t}_1$  the  $\tilde{g}$  decays to both  $b\tilde{b}_1$ and  $t\tilde{t}_1$ . For the scenarios represented by BMP-B1, BMP-B3 and BMP-C the BRs of the two modes are approximately the same irrespective of  $m_{\tilde{t}_1}$ . In BMP-B2 the BRs of the final states with  $\tilde{t}_1$  decreases as  $m_{\tilde{t}_1}$  approaches  $m_{\tilde{g}}$  and the scenario is indistinguishable to Gluino-sbottom-R model. We have chosen BMP-B2 such that  $m_{\tilde{t}_1}$  the  $\tilde{g}$  decays to both modes with approximately the same BR. As discussed above, in the limit of small L-R mixing  $\tilde{t}_1$  decays into a single channel channel with almost 100 % BR for  $m_{\tilde{t}_1}$  in a specific range. BMP-B1 and BMP-B2 represent scenarios with qualitatively different stop decays and can in principle be distinguished from other scenarios. On the other hand BMP-B3 and BMP-C are indistinguishable so far as stop decays are concerned.

We ensure that the spectrum, generated by SuSpect [74], contains a CP-even (light) Higgs with mass  $125 \pm 3 \text{ GeV}$ , as required by the present LHC data [75–78]. The relevant modes for  $\tilde{g}$ ,  $\tilde{b}_1$  and  $\tilde{t}_1$  decays with their branching ratios (BRs) for the benchmark points are shown in table 2. We have used SUSYHIT [79] to obtain the decay widths and branching ratios in various channels. We have computed the DM relic density using micrOMEGAs-3.6.7. [80].

For the Gluino-Sbottom-R-Model, in estimating the thermal abundance, three processes play major roles. The first one is  $\tilde{b}_1 \tilde{b}_1^* \to gg$ , which contributes dominantly to  $\sigma_{\text{eff}}$ . This process receives contributions from four-point (gauge) interaction, s-channel g exchange, and t/u-channel  $\tilde{b}$  exchange processes where the first two channels contribute most. In addition, there can be annihilation via  $\tilde{\chi}_1^0 \tilde{b}_1 \to gb$  and  $\tilde{b}_1 \tilde{b}_1 \to bb$ . Note that  $\tilde{\chi}_1^0$  annihilation is a pure EW process and hence its contribution remains small. While s-channel b exchange and t/u-channel  $\tilde{b}_1$  exchange contributes to the former one; the latter one is mediated by  $\tilde{g}$  and neutralinos (mostly the bino-like one) in t/u-channel. For our BMP-A, these three processes contribute 63%, 18% and 11% respectively. Although, because of large  $m_{\tilde{g}}$  mass, the contribution from the third process is rather small for smaller  $\tilde{\chi}_1^0$  masses, it increases for larger  $\tilde{\chi}_1^0$  masses contributing about 20% for  $m_{\tilde{\chi}_1^0} \sim 600 \,\text{GeV}$ .

In the Gluino-Stop-Sbottom-R-Model, for BMP-B1,  $\tilde{b}_1\tilde{b}_1^* \to gg$  contributes about 33%;  $\tilde{t}_1\tilde{b}_1 \to gb$  and  $\tilde{\chi}_1^0\tilde{t}_1 \to gt$  contribute about 15% each; and  $\tilde{\chi}_1^0\tilde{\chi}_1^0 \to t\bar{t}$  contributes about 11%. The latter receives contributions from  $\tilde{t}_1$  exchange t/u-channel processes, and is unsuppressed due to the large top mass which helps in evading large chirality suppression. Any other channel contributes less than 10% for this chosen benchmark point. In BMP-B2 and BMP-B3,  $\tilde{t}_1$  is quite heavy compared to  $\tilde{\chi}_1^0$ . Consequently, the leading co-annihilation channels involve  $\tilde{b}_1$  only, and their relative contributions are similar to that in Gluino-Sbottom-R-Model (BMP-A). However, since  $\tilde{t}_1$  is quite light compared to that in BMP-A,  $\tilde{\chi}_1^0$  annihilation into  $t\bar{t}$  (mediated via  $\tilde{t}_1$  in t/u-channel) contributes about 5 - 6% in both these cases.

Finally, in the *Gluino-Stop-Sbottom-L-Model*, the leading contribution comes from  $\tilde{b}_1 \tilde{b}_1^* \to gg$ , which gives about 34% of the total coannihilation rate for BMP-C. Since  $\tilde{b}_1$  is L-type, large D-term contribution together with contributions from the F-term, leads to large  $\tilde{b}_1 \tilde{b}_1^* \to hh$  cross-section. Along with four-point vertices, *s*-channel *h* mediated process and t/u-channel  $\tilde{b}_1$  mediated processes contribute to this channel. Altogether, its contribution is about 22%. This is followed by the cross-sections for  $\tilde{b}_1 \tilde{b}_1 \to bb$ , and other channels which contribute less than 10% for BMP-C. Since  $m_{\tilde{t}_1}$  is quite large (due to large D-term and F-term contribution, in the no-mixing limit), co-annihilation with  $\tilde{t}_1$  does not take place.

In figure 1, the red points show the region where  $\tilde{\chi}_1^0$  has the right thermal relic abundance in *Gluino-Sbottom-R-Model*. The mass difference with  $\tilde{b}_1$  does not change significantly over a large region of  $m_{\tilde{\chi}_1^0}$ ; it varies between about 25–35 GeV. Since strong interaction processes involving  $\tilde{b}_1$  always contribute dominantly to the relic density, such a minor variation of  $(m_{\tilde{b}_1} - m_{\tilde{\chi}_1^0})$  is well anticipated, especially since  $\sigma_{\text{eff}}$  depends exponentially on the same. The figure further shows the present ATLAS limit on  $m_{\tilde{b}_1}$  derived from  $0l + 2b + \not{E}_T$  and  $0l + 1j + \not{E}_T$  (compressed scenario) for direct  $\tilde{b}_1$  production [58, 59].

In *Gluino-Stop-Sbottom-R-Model*  $\tilde{t}_1$  (*R*-type) also contributes to the co-annihilation processes; for BMP-B1 its contribution is about 30%. Correspondingly, the mass difference  $(m_{\tilde{b}_1} - m_{\tilde{\chi}_1^0})$  is raised by about 4 GeV compared to that in *Gluino-Sbottom-R-Model* scenario. The mass difference varies between about 28-35 GeV over a large range (up to about 800 GeV) of  $m_{\tilde{\chi}_1^0}$  in this scenario too.

Green (blue) Points in figure 2a(2b) represent the parameter space with correct thermal relic density allowed by PLANCK in the Gluino-Stop-Sbottom-R(L)-Model. Orange and pink shaded regions are excluded from direct  $\tilde{b}_1$  pair production data in  $0l + 2b + \not{\!\!E}_T$ channel [58] and  $0l + \text{monojet} + \not{\!\!E}_T$  channel [59] by ATLAS.

Note that, we have used gluino masses around 1200 GeV to estimate the relic density. However, the t (and u) channel gluino exchange processes  $\tilde{b}_1 \tilde{b}_1 \rightarrow b b$  contributes only about 10-25% to  $\sigma_{\text{eff}}$  during DM freeze-out for our benchmark points.<sup>2</sup> Even a large increment

<sup>&</sup>lt;sup>2</sup> In  $\tilde{b}_1 \tilde{b}_1^* \to b\bar{b}$  s channel g exchange contributes dominantly along with t (and u) channel  $\tilde{g}$  exchange



in  $m_{\widetilde{g}}$ , therefore, can be compensated by a corresponding reduction in the mass of the NLSP by a few GeV, retaining the correct thermal relic abundance. We have checked this numerically using micrOMEGAs [80].<sup>3</sup>

It should be further noted that we have assumed  $|\mu| = 1000$ GeV. While reducing  $|\mu|$  has no effect on the (2-body) decay of  $\tilde{g}$ ; significant bino-higgsino mixing in  $\tilde{\chi}_1^0$  can be achieved. However, such mixing is quite constrained by the present data from LUX [81]. For a 300 GeV  $\tilde{\chi}_1^0$ , we found it to be at most about 5-10%. But the details of direct detection cross-section also depend on  $\tan \beta$  and the heavy Higgs masses. Moreover, uncertainties from the estimation of the strange quark content of the nucleon affect the direct detection cross-section significantly; assuming the default values of strange quark content at low energies may turn out to be an oversimplification [80]. In the present context, its effect on the thermally averaged annihilation cross-section can again be compensated by changing the mass of the NLSP by a few GeV, thus restoring the correct thermal relic abundance. It may thus be worthwhile to consider sbottom NLSP scenarios with low values of  $|\mu|$  as well.

In a nutshell, we have considered various scenarios with a light  $\tilde{b}_1$  NLSP (and also light  $\tilde{t}_1$  in benchmarks BMP-B and BMP-C). In all these scenarios, the bino-like  $\tilde{\chi}_1^0$  provides a suitable Dark Matter candidate. While its dominantly bino-like nature assures that it

process. However, this channel contributes less than 1% to the thermally averaged annihilation cross-section.

<sup>&</sup>lt;sup>3</sup>Note that, the mass difference between the LSP and the NLSP affects the contributions from the co-annihilating channels to  $\sigma_{\text{eff}}$  exponentially.



Figure 2: The green and blue points give correct thermal relic density of  $\tilde{\chi}_1^0$  allowed by PLANCK [69] in Gluino-Stop-Sbottom-R-Model (left panel) and Gluino-Stop-Sbottom-L-Model (right panel) respectively (see text for details). Limits on  $m_{\tilde{b}_1}$  from direct  $\tilde{b}_1$  pair production from  $0l + \text{monojet} + \not{\!\!\!E}_T$  search [59] and  $0l + 2b + \not{\!\!\!E}_T$  search [58] by ATLAS are presented by orange and pink shaded region respectively.

escapes the present direct searches, co-annihilation with  $\tilde{b}_1$  plays a crucial role in achieving the right thermal relic density. In the next section, we will consider the lower bound on the mass of  $\tilde{g}$  in these scenarios.

### 3 Constraints on the $\tilde{b}_1 - \tilde{g}$ sector in the $\tilde{b}_1$ NLSP scenarios

In this section we focus on  $\tilde{g}$  pair production in the sbottom NLSP scenarios described in the previous section and obtain the exclusion contours in the  $m_{\tilde{g}} - m_{\tilde{b}_1}$  plane using the LHC data. For each point in the plot  $m_{\tilde{\chi}_1^0}$  is chosen in the narrow range consistent with the observed DM relic density. We also compute the limits on  $m_{\tilde{g}}$  for different models described in previous section and also for different BMPs introduced in table 1 and 2. Since there is no dedicated LHC search for this case our aim is to constrain the gluino mass utilizing the ATLAS model independent cross section bounds (see below) for final states similar to but not identical with the ones we are interested in. Our choice is guided by the fact that the final state coming from  $\tilde{g}\tilde{g}$  pair production in all cases is expected to be enriched with *b*-jets and the generic signal will be jets +  $\not{E}_T$ .

in most cases of our interest, it yields the most stringent bounds for significantly smaller values of  $\Delta m_{\tilde{b}_1}$ . Such choices, though disfavoured in the  $\tilde{b}_1$  -  $\tilde{\chi}_1^0$  co- annihilation scenarios, can not be absolutely ruled out in view of the fact that DM production may be due to other mechanisms.

Next we will briefly review the above analyses by ATLAS. SUSY searches in the inclusive jets +  $0l + \not\!\!\!E_T$  channel for  $\mathcal{L} = 20.3 \text{ fb}^{-1}$  at 8 TeV have been presented in ref. [12]. Five inclusive analyses channels, labelled as A to E depending on jet multiplicities (from two to six respectively), are introduced. The relevant cuts have been described in table 1 of ref. [12]. Each of these channels are further classified as 'Tight', 'Medium' and 'Loose' based on the final cuts on the observables  $\frac{\not\!\!\!E_T}{m_{\text{eff}}}$  and  $m_{\text{eff}}$ (incl.).<sup>4</sup> Non-observation of any significant excess in each of these signal region leads to an upper limit on the number of events ( $N_{BSM}$ )from any Beyond Standard Model (BSM) scenario. The observed upper limits on  $N_{BSM}$  at 95 % Confidence Level (CL) for signal regions SRA-Light, SRA-Medium, SRB-Medium, SRB-Tight, SRC-Medium, SRC-Tight, SRD, SRC-Loose, SRE-Tight, SRE-Medium, SRE-Loose are given by 1341, 51.3, 14.9, 6.7, 81.2, 2.4, 15.5, 92.4, 28.6, 8.3 respectively [12]. We use these (model independent) numbers to obtain new limits on  $m_{\widetilde{a}}$ .

ATLAS collaboration has also reported the search for strong sparticles in the multib-jets final states with  $\mathcal{L} = 20.1 \text{ fb}^{-1}$  at 8 TeV in [13]. Selection criteria for the signal regions are listed in table 1 and 2 of ref. [13]. In this analysis both 0l (two signal regions) and at least one lepton (one signal region) channel are introduced. Signal regions are characterized by large  $\not{E}_T$  and at least four (SR - 0l - 4j), six (SR - 1l - 6j) or seven (SR - 0l - 7j) jets which includes at least three b-tagged jets. Finally, these are further classified as A/B/C depending on  $\not{E}_T$  and  $m_{\text{eff}}$ . The absence of any excess led to an upper limits on the number of signal events in each of these regions from ATLAS. In particular, for signal regions SR-0l-4j-A, SR-0l-4j-B, SR-0l-4j-C, SR-0l-7j-A, SR-0l-7j-B, SR-0l-7j-C, SR-1l-6j-A, SR-1l-6j-B, SR-1l-6j-C these upper limits, at 95 % CL, are given by 4.6, 6.7, 4.8, 15.3, 6.1, 3.9, 6.6, 3.0, 3.0 respectively.

We adopt different selection criteria for various signal regions discussed above. For electron, muon and jet identification, reconstruction, isolation etc, we use the ATLAS prescription described in the respective analyses [12, 13].<sup>5</sup> The  $P_T$  dependent b-tagging efficiencies provided by ATLAS collaboration in ref. [82] have been used in our code. For validation, we reproduced the number of events in each of these signal regions, as obtained by ATLAS, for some benchmark points in refs. [12–15].

We then use PYTHIA (v6.428) [83] to generate the signal events in different channels from gluino pair production for different scenarios. The NLO cross-section for the  $\tilde{g}\tilde{g}$  pair production is computed with PROSPINO 2.1 [84] using CTEQ6.6M PDF [85]. Finally we derive the new limits on gluino mass by comparing the computed number of events with the

 $<sup>{}^4</sup>m_{\rm eff}$  is defined as the scalar sum of the transverse momenta of the leading N jets which defines the signal region and  $\not\!\!\!E_T$ .  $m_{\rm eff}$  (incl.) is defined as the scalar sum of the transverse momenta of the jets having  $P_T$  greater than 40 GeV and  $\not\!\!\!E_T$ .



**Figure 3**: Limits on  $m_{\widetilde{g}}$  in Gluino-Sbottom-R-Model from  $\widetilde{g}\widetilde{g}$  pair production in  $m_{\widetilde{g}} - m_{\widetilde{b}_1}$ plane from  $0l + jets + \not{E}_T$  [12] and  $0l + jets(3b) + \not{E}_T$  channel [13] at 95 % *CL* with 8 TeV 20 fb<sup>-1</sup> ATLAS data. The shaded region are excluded from direct  $\widetilde{b}_1$  pair production [58, 59] (see figure 1)

corresponding upper limits on  $N_{BSM}$  in the relevant signal regions. The exclusion regions for Gluino-Sbottom-R-Model, thus obtained, are presented in figure 3.

In figure 3 each point gives correct DM relic density via sbottom coannihilation i.e., mass difference between  $\tilde{b}_1$  and  $\tilde{\chi}_1^0$  is in the range 25–35 GeV. We only consider the parameter space  $m_{\tilde{g}} > m_{\tilde{b}_1} + 100$  GeV. If  $m_{\tilde{g}}$  is smaller then, as discussed in section 2, the gluino contributes to DM relic production significantly leading to under abundant relic density. The shaded region in figure 3 is excluded by direct search for sbottom pairs for low  $\Delta m_{\tilde{b}_1}$ .

The most stringent bound on  $m_{\widetilde{g}}$  comes from  $0l + jets(3b) + \not{E}_T$  channel [13] — for  $m_{\widetilde{b}_1}$  around 500 GeV, gluino mass below 1.1–1.2 TeV is excluded. The red and blue lines in figure 3 represents the exclusion contour from  $0l + jets + \not{E}_T$  [12] and  $0l + jets(3b) + \not{E}_T$  channel [13]. Due to relatively soft b-jets from  $\tilde{b}_1$  decays, the signal in the sbottom NLSP scenario is sensitive to signal regions with low jet multiplicity. The two most effective signal regions are SRB-Medium [12] and SR-0l-4j-A [13] which require 3jets and 4jets respectively. In the Gluino-Stop-Sbottom-R-Model depending upon stop decay modes we consider three possibilities (see and table 1 and table 2). If the decay  $\tilde{t}_1 \to c \tilde{\chi}_1^0$  is the most dominant mode (see BMP-B1), then the gluino mass limits is expected to be rather weak. For our computations in this case we assume the mass relations  $m_{\tilde{b}_1} = m_{\tilde{\chi}_1^0} + 40$  GeV. As already noted the relatively light  $\tilde{t}_1$  also contributes to DM production. We present the results in figure 4 for this model. Gluino limits in this model



Figure 4: Limits on  $m_{\widetilde{g}}$  in Gluino-Stop-Sbottom-R-Model from  $\widetilde{g}\widetilde{g}$  pair production in  $m_{\widetilde{g}} - m_{\widetilde{b}_1}$  plane from  $0l + jets + \not{E}_T$  [12] and  $0l + jets(3b) + \not{E}_T$  channel [13] at 95 % CL with 8 TeV 20 fb<sup>-1</sup> ATLAS data. Here  $BR(\widetilde{b}_1 \to b\widetilde{\chi}_1^0) = 100\%$  and  $\widetilde{t}_1 \to c\widetilde{\chi}_1^0$  dominates (for illustration see BMP-B1). The shaded region are excluded from direct  $\widetilde{b}_1$  pair production [58, 59] (see figure 1)

are weaker than in figure 3 roughly by 100 GeV in both channels. This is due to the fact that the gluino decays into  $\tilde{t}_1 t$  with almost 50% BRs and effectively reduce the 0l or bjet signal events. It may be noted that (see figure 4) when  $\tilde{g} \to \tilde{t}_1 t$  is kinematically disfavoured, then the limits tend to increase to some extent.

In the other scenarios where  $\tilde{t}_1 \to t \tilde{\chi}_1^0$  or  $\tilde{t}_1 \to W \tilde{b}_1$  dominates, there are extra sources of jets and tagged bjets. As a result the limits on  $m_{\tilde{g}}$  are pushed up. As the exclusion contour or mass limits lie between the results obtained in figure 3 and figure 4, we have not presented any detailed figure for such scenarios. Instead we present the limits on  $m_{\tilde{g}}$ for the benchamark points (see table 3). As expected BMP-A gives the strongest limit  $m_{\tilde{g}} \gtrsim 1.2 \text{ TeV}$  and for BMP-B1 it reduced to 1.1 TeV. We note in passing that for BMP-B2, BMP-B3 and BMP-C the signal contain multiple Ws which reduces the limits from the  $0l + jets + \not{\!\!\!E}_T$  channel [12]. In such cases, however, the hard single lepton channel  $(1l + jets + \not{\!\!E}_T)$  [14] puts more stringent bounds ( $m_{\tilde{a}} > 1025-1050 \text{ GeV}$ ).

Points	Limit on $m_{\widetilde{q}}$ (GeV)			
	$0l + jets + \not\!\!\!E_T$ [12]	$0l + jets(3b) + \not\!$		
BMP-A	1070	1195		
BMP-B1	1000	1115		
BMP-B2	845	1165		
BMP-B3	850	1135		
BMP-C	850	1150		

**Table 3**: Limits on  $m_{\widetilde{g}}$  using the ATLAS  $0l + jets + \not\!\!\!E_T$  data [12],  $0l + jets(3b) + \not\!\!\!E_T$  data [13] for the benchmarks.

In the  $\tilde{t}_1$  NLSP scenarios with correct DM relic density, the decays  $\tilde{b}_1 W$  and  $t \tilde{\chi}_1^0$  are not allowed. However there are sufficiently large parameter space in the  $\tilde{b}_1$  NLSP scenario where these decays are allowed. These decays may provide novel signatures of this scenario during the LHC phase 2 experiments.

### 4 Conclusion

In view of the current interest in SUSY with light third generation squarks, we have studied the co-annihilating sbottom NLSP scenario and have delineated the parameter space consistent with the DM relic density constraint (see figure 1) when the NLSP is of R-type with both of the stop mass eigenstates being rather heavy (the Gluino-Sbottom-R model). It easier to fit the observed Higgs mass in this case. It should also be stressed that in this scenario the co-annihilation cross section is relatively large since strong interactions partially contribute to it. As a result somewhat large (25–35 GeV)  $\Delta m_{\tilde{b}_1}$  is consistent with the PLANCK data (see eq. 2.1). This enhances the observability of the LHC signatures.

However, there are other options with an additional light  $t_1$  like the Gluino-Sbottom-Stop-R model and the Gluino-Stop-Sbottom-L-Model (section 2). The allowed range of  $\Delta m_{\tilde{b}_1}$  is very similar to the one in the previous model (see figure 2). Some of the models predict spectacular LHC signatures with multiple tagged b-jets, W's and reconstructable t's over a large parameter space which should be searched for during phase 2 (section 2). These signals are not viable in the  $\tilde{t}_1$ -NLSP scenario. The fraction of events with reconstructed W, top etc. can reveal the underlying scenarios including some hints on the mixing angles in the stop and sbottom sectors.

We note in passing that the  $(g-2)_{\mu}$  anomaly [86, 87] can be easily resolved within the framework of the  $\tilde{b}_1$  NLSP scenario with relatively heavy electroweak sparticles. For example, if the sleptons are heavy,  $m_{\tilde{\chi}_1^0} \gtrsim 100 \text{ GeV}$ , the constraint on the lighter chargino  $(\tilde{\chi}_1^{\pm})$  mass from LHC is very weak and a large parameter space is consistent with the  $(g-2)_{\mu}$ constraint. For lighter LSP there are stronger LHC constraints on the chargino mass; yet a reasonable parameter space compatible with the  $(g-2)_{\mu}$  constraint is allowed (for a recent discussion see figure 8 and section 3.4 of ref. [73]). In the Gluino-Sbotom-R scenario the allowed  $\tilde{b}_1$ -NLSP masses are bounded from below by the LHC lower limit (260 GeV) from direct searches of coannihilating sbottom squarks. On the other hand  $m_{\tilde{b}_1}$  is bounded from above by the maximum chargino mass determined by the  $(g-2)_{\mu}$  constraint, the LSP mass and the LHC searches for the electroweak sparticles. The same conclusion holds for Gluino-Sbottom-Stop-R model with  $m_{\tilde{b}_1}$  and  $m_{\tilde{t}_1}$  close together. In the variants of the  $\tilde{b}_1$ NLSP scenario with a heavier  $\tilde{t}_1$  the mode  $\tilde{t}_1 \rightarrow b \tilde{\chi}_1^{\pm}$  may open up.

There is no dedicated search for the  $\tilde{b}_1$  NLSP production from gluino decays as yet. We obtain new constraints on the  $\tilde{g} - \tilde{b}_1$  sector in all scenarios using the existing LHC searches involving somewhat similar final states. We find that for  $m_{sbot}$  around 500 GeV the limits on  $m_{\tilde{g}}$  in the  $0l + jets(3b) + \not{E}_T$  [13] channel vary in the range 1.1-1.2 TeV in different scenarios (see figure 3, 4 and table 3). For higher  $m_{sbot}$  the limit is degraded rapidly. On the other hand, for small  $\Delta m_{\tilde{b}_1}$  and  $m_{\tilde{g}} \approx 1 \, TeV$ , much stronger limits on the  $\tilde{b}_1$  mass is obtained via  $\tilde{g}\tilde{g}$  pair production compared to that obtained from direct  $\tilde{b}_1$  searches. We find that the  $0l + jets + \not{E}_T$  data generally yield somewhat weaker bounds. However, if the relic density constraint is relaxed and  $\Delta m_{\tilde{b}_1}$  is allowed to be smaller ( $\approx 10 \, \text{GeV}$ ) the strongest limit ( $\approx 1 \, \text{TeV}$ ) is obtained from the  $0l + jets + \not{E}_T$  data. On the whole the gluino mass limit in the light sbottom scenario is about 1 TeV irrespective of  $\Delta m_{\tilde{b}_1}$ .

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